

Nonlinear Schrödinger equation under a magnetic field kept in the foreground

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Nonlinear Schrödinger equation when the magnetic field is kept in the foreground

- 1 The magnetic nonlinear Schrödinger equation
- 2 Semiclassical results
- 3 Constant magnetic field problem

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Magnetic nonlinear Schrödinger functional

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where

- $u : \Omega \rightarrow \mathbb{C} \simeq \mathbb{R}^2$ is a wave-function,
- $-|u|^{p-2}$ is an interaction by δ -functions,
- V is an (adimensionalized) electric potential,
- ε is a characteristic length (adimensionalized Planck constant),
- $D_{A/\varepsilon}$ is a covariant derivative.

Covariant derivative and connection

Differential geometry meets physics

$$D_A u \triangleq Du - iuA,$$

where $A : \Omega \rightarrow \wedge^1 \mathbb{R}^3 \simeq \mathbb{R}^3$.

① differential calculus:

$$D_A(\eta u) = \eta D_A u + u D\eta,$$

$$D|u|^2 = 2(u|D_A u|),$$

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② the curvature is the magnetic field $B = dA \simeq \nabla \times A$,

$$\begin{aligned} K_A[v, w]u &\triangleq D_A(D_A u[w])[v] - D_A(D_A u[v])[w] \\ &= -idA[v, w]u \simeq -i(\nabla \times A|v \times w)u, \end{aligned}$$

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3 gauge invariance: if $A' \triangleq A + d\lambda \simeq A + \nabla\lambda$ and $u' \triangleq e^{i\lambda}u$, then

$$D_{A'} u' = e^{i\lambda} D_A u.$$

The Euler–Lagrange equation associated to

$$\int_{\Omega} \frac{\varepsilon^2 |D_{A/\varepsilon} u|^2}{2} + \frac{V|u|^2}{2} - \frac{|u|^p}{p}$$

is

$$-\varepsilon^2 \Delta_{A/\varepsilon} u + Vu = |u|^{p-2} u,$$

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where the **magnetic Laplacian** is defined as

$$\begin{aligned} -\varepsilon^2 \Delta_{A/\varepsilon} u &= -\varepsilon^2 \Delta u + 2i\varepsilon \langle A, Du \rangle + (|A|^2 - id^* A)u \\ &\simeq -\varepsilon^2 \Delta u + 2i\varepsilon A \cdot \nabla u + (|A|^2 - i\nabla \cdot A)u. \end{aligned}$$

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in the **Coulomb/Lorenz gauge** ($d^* A \simeq \nabla \cdot A = 0$).

Existence of groundstates

Seminal work of Esteban and Lions

Groundstate solutions are minimizers of

$$\frac{\int_{\Omega} \varepsilon^2 |D_{A/\varepsilon} u|^2 + V|u|^2}{\left(\int_{\Omega} |u|^p\right)^{\frac{2}{p}}}.$$

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Theorem (M. Esteban and P.-L. Lions, 1989)

Groundstates exist when either

- 1 Ω is bounded,
- 2 $\Omega = \mathbb{R}^3$, and V and dA are *constant*,
- 3 a strict inequality is satisfied.

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Magnetic field going into the background

Magnetic field in the background

Theorem (Kurata (2000))

If $\frac{1}{2} - \frac{1}{3} < \frac{1}{p} < \frac{1}{2}$ and $V_0 \triangleq \inf V < \liminf_{\infty} V$, then there exist $(x_\varepsilon)_{\varepsilon>0}$ in \mathbb{R}^3 and $(\omega_\varepsilon)_{\varepsilon>0}$ such that $V(x_\varepsilon) \rightarrow V_0$ and

$$u_\varepsilon(x) \simeq \alpha_\varepsilon U\left(\frac{x - x_\varepsilon}{\varepsilon}\right) e^{i\left(\frac{A(x_\varepsilon)[x - x_\varepsilon]}{\varepsilon} + \omega_\varepsilon\right)},$$

where U is the unique positive radial groundstate of the limiting problem

$$-\Delta U + V_0 U = |U|^{p-2} U.$$

See also Cingolani (2003), Cingolani & Secchi (2002), Secchi & Squassina (2005), Cao & Tang (2006), Barile (2008), Cingolani, Jeanjean & Secchi (2009), Cingolani & Clapp (2009, 2010), Barile, Cingolani & Secchi (2011), Ding & Liu (2013)...

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The behavior of the concentration points and of the densities are independent of the magnetic field:

$$|u_\varepsilon(x)|^2 \simeq \left| U\left(\frac{x - x_\varepsilon}{\varepsilon}\right) \right|^2.$$

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Asymptotically consistent with the Lorentz force when $v = 0$:

$$0 = F = -q(dV + v \lrcorner dA) \simeq q(-\nabla V + v \times (\nabla \times A)).$$

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Question: How to have interaction with the magnetic field in the semiclassical limit?

Magnetic semiclassical soliton dynamics

The Lorentz force is validated in the dynamics

Theorem (Selvitella (2008) & Squassina (2009))

If $2 < p < 2 + \frac{4}{3}$, then there exists solutions to

$$i\varepsilon\partial_t\psi = -\varepsilon^2\Delta_{A/\varepsilon}\psi + V\psi - |\psi|^{p-2}\psi,$$

of the form

$$\psi_\varepsilon(x, t) \simeq U\left(\frac{x - x(t)}{\varepsilon}\right) e^{i\left(\frac{A(x(t))[x-x(t)] + \dot{x}(t)\cdot(x-x(t))}{\varepsilon} + \omega_\varepsilon(t)\right)}.$$

with

$$\begin{aligned}\ddot{x}(t) &= -dV(x(t)) - \dot{x}(t) \lrcorner dA(x(t)) \\ &\simeq -\nabla V(x(t)) + \dot{x}(t) \times (\nabla \times A)(\dot{x}(t)).\end{aligned}$$

Why is the magnetic field fading out?

By the Heisenberg uncertainty principle, a groundstate might have a **magnetic momentum** $\mu \in \bigwedge^2 \mathbb{R}^3 \simeq \mathbb{R}^3$. It is thus subject to a force of

$$-qdV + d\langle dA, \mu \rangle \simeq -q\nabla V + \nabla(\nabla \times A \cdot \mu).$$

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There is **no asymptotic magnetic effect**:

$$q = \int_{\mathbb{R}^3} |\nabla u_\varepsilon|^2 \simeq \varepsilon^3, \quad \mu = \int_{\mathbb{R}^3} ((x - x_\varepsilon) \wedge (-i\varepsilon \nabla_{A/\varepsilon} u_\varepsilon(x)) |u_\varepsilon(x)|) dx \simeq \varepsilon^4.$$

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Idea: Interesting phenomena should happen in the **strong magnetic field régime** $A \simeq \varepsilon^{-1}$.

The new problem

We study solutions of the Euler–Lagrange equation associated to

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which is

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The concentration function

For $V \in \mathbb{R}$ and $B \in \bigwedge^2 \mathbb{R}^3 \simeq \mathbb{R}^3$, define

$$A(x)[v] = \frac{B(x, v)}{2} \simeq \frac{(B \times x) \cdot v}{2},$$

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$$\mathcal{E}(V, B) = \left(\frac{1}{2} - \frac{1}{p} \right) \inf_{u \in C_c^\infty(\mathbb{R}^3; \mathbb{C})} \frac{\left(\int_{\mathbb{R}^3} |D_A u|^2 + V|u|^2 \right)^{\frac{p}{p-2}}}{\left(\int_{\mathbb{R}^3} |u|^p \right)^{\frac{2}{p-2}}}.$$

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- 1 \mathcal{E} is continuous,
- 2 diamagnetic inequality:

$$\mathcal{E}(V, B) \geq \mathcal{E}(V, 0).$$

Asymptotic result

Theorem (Di Cosmo and Van Schaftingen (2015))

If $\frac{1}{2} - \frac{1}{3} < \frac{1}{p} < \frac{1}{2}$ and Ω is bounded, then there exist $(x_\varepsilon)_{\varepsilon>0}$ in \mathbb{R}^3 such that

$$\lim_{\substack{R \rightarrow \infty \\ \varepsilon \rightarrow 0}} \|u_\varepsilon\|_{L^\infty(\Omega \setminus B_{\varepsilon R}(x_\varepsilon))} = 0,$$

$$\lim_{\varepsilon \rightarrow 0} \varepsilon^{-N} \mathcal{F}_\varepsilon(u_\varepsilon) = \lim_{\varepsilon \rightarrow 0} \mathcal{E}(V(x_\varepsilon), dA(x_\varepsilon)) = \inf_{\Omega} \mathcal{E}(V, dA).$$

If $V(x_{\varepsilon_n}) \rightarrow V_*$, $dA(x_n) \rightarrow B_*$ and $A_*(x)[v] = \frac{B_*(x,v)}{2}$, up to a subsequence

$$u_{\varepsilon_n} \simeq U_{V_*, A_*} \left(\frac{x - x_n}{\varepsilon_n} \right) e^{i \left(\frac{A(x_{\varepsilon_n})[x - x_{\varepsilon_n}]}{\varepsilon_n^2} + \omega_n \right)},$$

with U_{V_*, A_*} a groundstate of

$$-\Delta_{A_*} U_{V_*, A_*} + V_* U_{V_*, A_*} = |U_{V_*, A_*}|^{p-2} U_{V_*, A_*}.$$

Asymptotic result

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See also Fournais and Raymond (2015).

Forces in the semiclassical limit

If $x_* \in \Omega$ and

$$\mathcal{E}(V(x_*), dA(x_*)) = \inf_{\Omega} \mathcal{E}(V, dA),$$

and U_{V_*, A_*} is the corresponding groundstate, then

$$qdV(x_*) = d\langle dA, \mu \rangle(x_*),$$

with

$$q = \int_{\mathbb{R}^3} |U_{V_*, A_*}|^2$$

and

$$\mu = \int_{\mathbb{R}^3} (y \wedge (-i\nabla_A U_{V_*, A_*}(y)) |U_{V_*, A_*}(y)|) dy.$$

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About the proof

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Also covers local minimizers (del Pino & Felmer penalization), with slow decay at infinity

$$\lim_{|x| \rightarrow \infty} V(x)|x|^2 > 0.$$

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The constant electromagnetic potential problem

What happens to the groundstate of

$$-\Delta_A + Vu = |u|^{p-2}u$$

when $V = 1$ and $A(x)[v] = \frac{1}{2}B[x, v] \simeq \frac{1}{2}(B \times x) \cdot v$ for some $B \in \bigwedge^2 \mathbb{R}^3$?

- How does the solution depend on B ?

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- How does the solution depend on B ?
- Is the solution symmetric?
- How does the solution decay?

Free electron in a magnetic field

The first Landau level

Groundstate of the **linear** Schrödinger equation satisfy

$$-\Delta_A \psi = E \psi.$$

If $A(x)[v] = \frac{1}{2}B(x, v) \simeq \frac{1}{2}(B \times x) \cdot v$, then a groundstate is given by

$$\psi(x) = e^{-\frac{|B \times x|^2}{4|B|}},$$

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$$E = |B|.$$

Properties

- E increases with $|B|$,

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Properties

- E increases with $|B|$,
- ψ is axially symmetric with respect to B ,
- ψ has Gaussian decay perpendiculary to B .

Groundstates under a small constant magnetic field

Theorem (Bonheure, Nys and Van Schaftingen, 2015)

If $V \simeq 1$ and $|A|^2 \lesssim 1$ then

- 1 $U_{A,V}$ is real and radially symmetric,
- 2 $U_{A,V}(x) \lesssim e^{-\frac{|x \cdot B|^2}{4|B|^2}} \simeq e^{-\frac{|B \times x|^2}{4|B|^2}}$,
- 3 $\mathcal{E}(V, B) \simeq \mathcal{E}(1, 0) + \left(\frac{V-1}{2} \int_{\mathbb{R}^3} |U|^2 + \frac{|B|^2}{4 \cdot 3} \int_{\mathbb{R}^3} |y|^2 |U(y)|^2 dy \right)$.

“Essentially” means up to translations in \mathbb{R}^3 and rotation in \mathbb{C} .

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By scaling, we cover in fact $|A|^2/V \lesssim 1$.

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In \mathbb{R}^2 , the Gaussian decay follows from the linear theory (L. Erdős, 1996).

Proving the properties by essential uniqueness

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Problem: U_A lives naturally in the space

$$H_A^1(\mathbb{R}^3) = \{u \in L^2(\mathbb{R}^3) : D_A u \in L^2(\mathbb{R}^3)\};$$

the natural functional space depends on the parameter A .

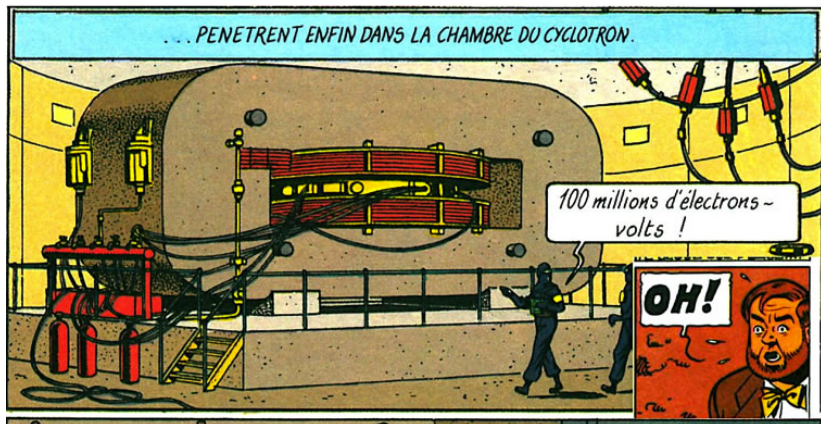
- the gap between $U_1 \in H_{A_1}^1(\mathbb{R}^3)$ and $U_2 \in H_{A_2}^1(\mathbb{R}^3)$ can be measured by

$$\int_{\mathbb{R}^3} |D_{A_1} U_1 - D_{A_2} U_2|^2 + |U_1 - U_2|^2,$$

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- the uniqueness part of the implicit function theorem just uses nondegeneracy and continuity of the linearized operator.



Thank you for your attention.